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CERN

Orientifolds

model building

extra dimensions

Hierarchy problem: why gravity is so weak compared to the other interactions?

Quantum theory: all particle masses $\nearrow M_P \sim 10^{19}$ GeV

- Supersymmetry: protection of hierarchy due to cancellations between fermions and bosons

$$\Rightarrow m_{\text{susy}} \sim \text{TeV}$$

- TeV strings: low UV cutoff

$$\Rightarrow M_s \sim \text{TeV}$$

- Split supersymmetry: unknown solution live with the hierarchy

$$\Rightarrow m_0 \text{ heavy, fermions light}$$

→ all of them testable at LHC

- Heterotic string:

Natural framework for susy and unification

However mismatch between string and GUT scales

$$M_s = gM_P \simeq 50M_{\text{GUT}}$$

- Framework of type I string theory

⇒ D-brane world

Natural separation of
global SUSY from gravity



D-branes/open strings

closed strings

⇒ 2 new scenaria besides 'conventional'

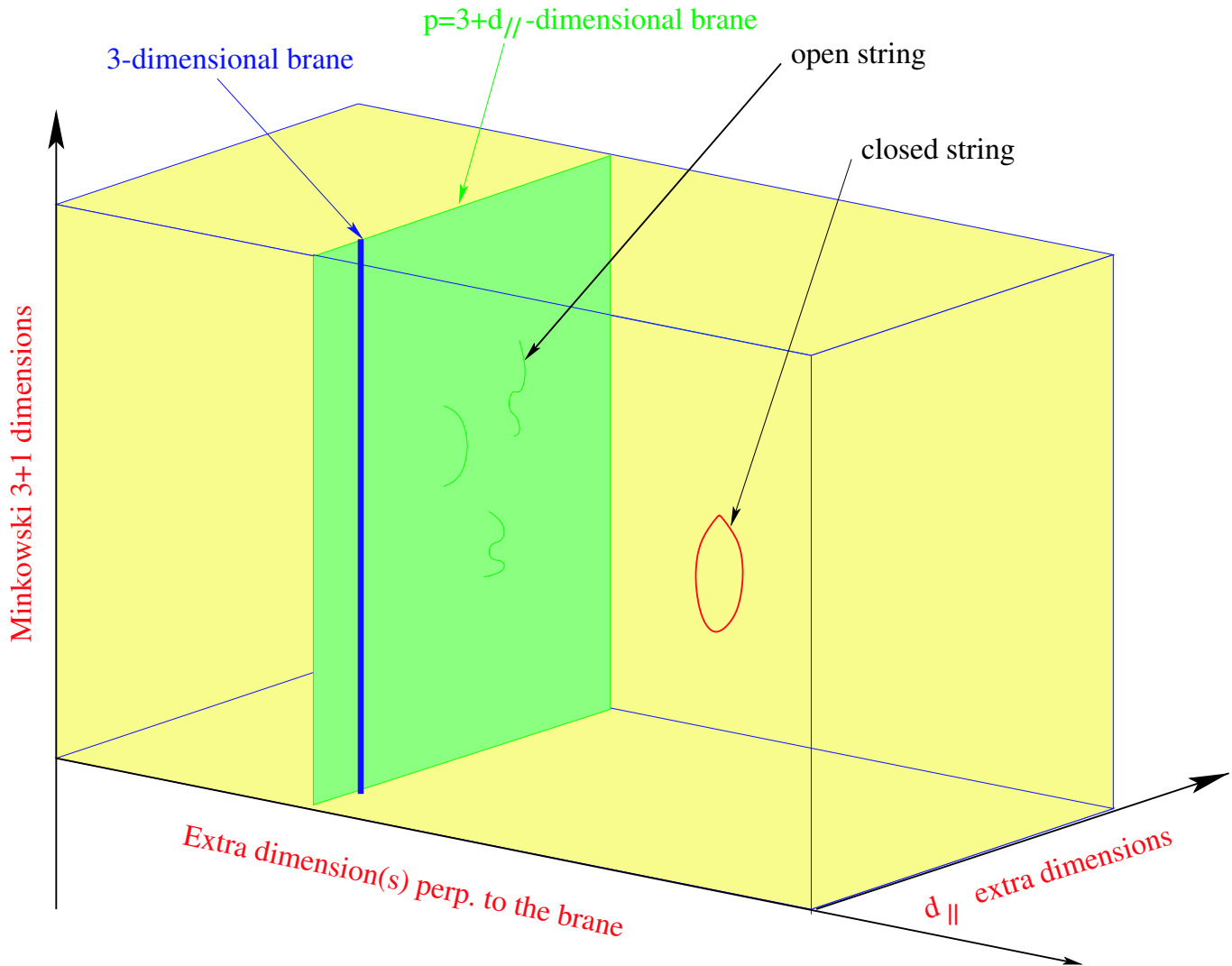
low energy susy Standard Model

- low string scale
- split supersymmetry

OUTLINE

- Framework of low scale strings
large extra dimensions, low scale gravity
- Experimental predictions
strong gravity, TeV dimensions, string effects
- SUSY in the bulk
brane SUSY breaking, short range forces
- Electroweak symmetry breaking
- D-brane embedding of the Standard Model
unification, proton stability, Right-neutrinos
- SUSY breaking by internal magnetic fields
or equivalently branes at angles
- Gaugino masses
Split supersymmetry, Dirac masses

Braneworld



two types of compact extra dimensions:

- parallel ($d_{||}$): can be as large as 10^{-16} cm (TeV^{-1})
- transverse (\perp): can be as large as 0.1 mm

I.A. '90

Dimensions of finite size: $p - 3$ parallel

$n = 9 - p$ transverse

calculability $\Rightarrow R_{\parallel} \simeq l_{\text{string}} ; R_{\perp}$ arbitrary

$$M_P^2 \simeq \frac{1}{\alpha^2} M_s^{2+n} R_{\perp}^n$$



Planck mass in $4 + n$ dims: M_*^{2+n}

small $M_s/M_P \Rightarrow$ extra-large R_{\perp}

$$M_s \sim 1 \text{ TeV} \Rightarrow R_{\perp} \sim .1 - 10^{-13} \text{ mm } (n = 2 - 6)$$

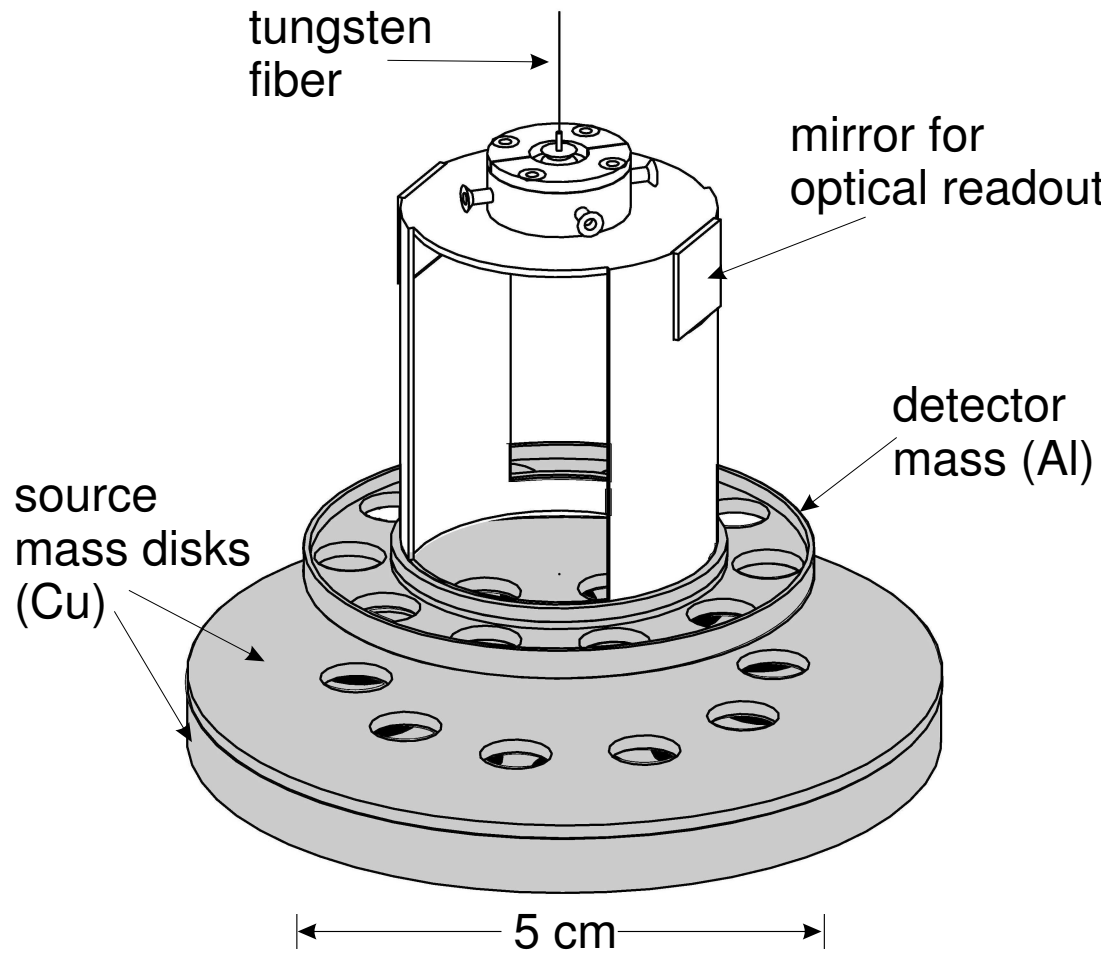
I.A.-Arkani Hamed-Dimopoulos-Dvali '98

- weak string coupling: $g_s = \alpha$
- gravity strong at $M_* \sim M_s \ll M_P$

10^{30} stronger than thought previously!

deviations from Newton's law at distances $< R_{\perp}$

Adelberger et al. '04



$R_{\perp} \lesssim 130 \mu\text{m}$ at 95% CL

Supernova constraints

cooling due to graviton production

e.g. $NN \rightarrow NN + \text{graviton}$

number of gravitons: $\sim (TR_{\perp})^n$ $T \gg R_{\perp}^{-1}$
 $\sim 10 \text{ MeV}$

\Rightarrow production rate:

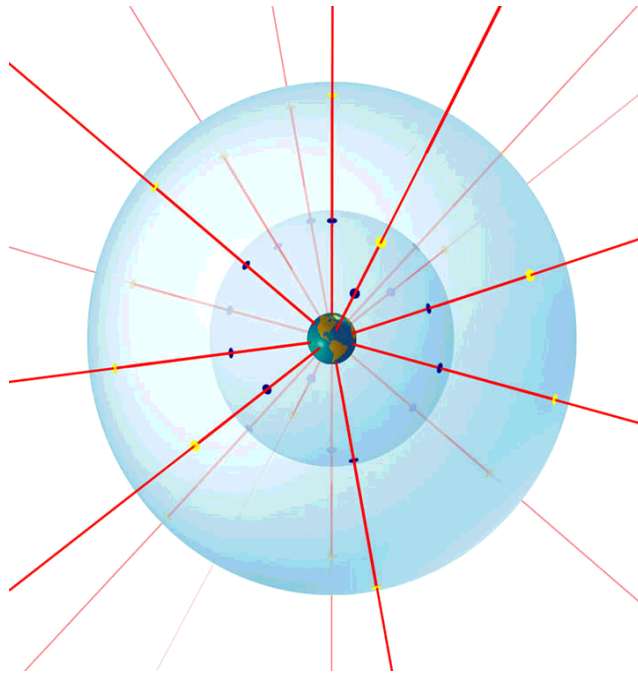
$$P_{\text{gr}} \sim \frac{1}{M_p^2} (TR_{\perp})^n \sim \frac{T^n}{M_*^{(2+n)}}$$

$$P_{\text{gr}} < P_{\nu} \Rightarrow M_*|_{n=2} \gtrsim 50 \text{ TeV}$$

$$\Rightarrow M_s \gtrsim 10 \text{ TeV}$$

Gravity modification at submillimeter distances

Newton's law: force decreases with area



3d: force $\sim 1/r^2$

$(3+n)$ d: force $\sim 1/r^{2+n}$

observable for $n = 2$: $1/r^4$ with $r \lesssim .1$ mm

Hidden submillimeter dimensions

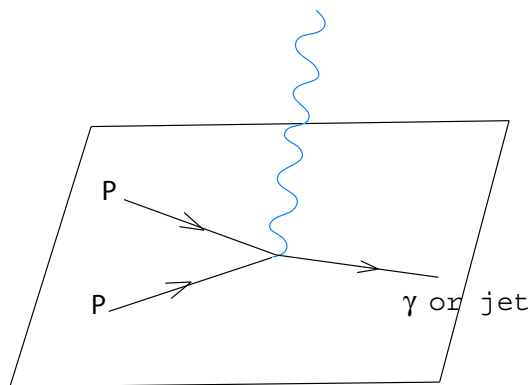
⇒ strong gravity at the TeV

Gravitational radiation in the bulk

3d: Kaluza Klein gravitons very light

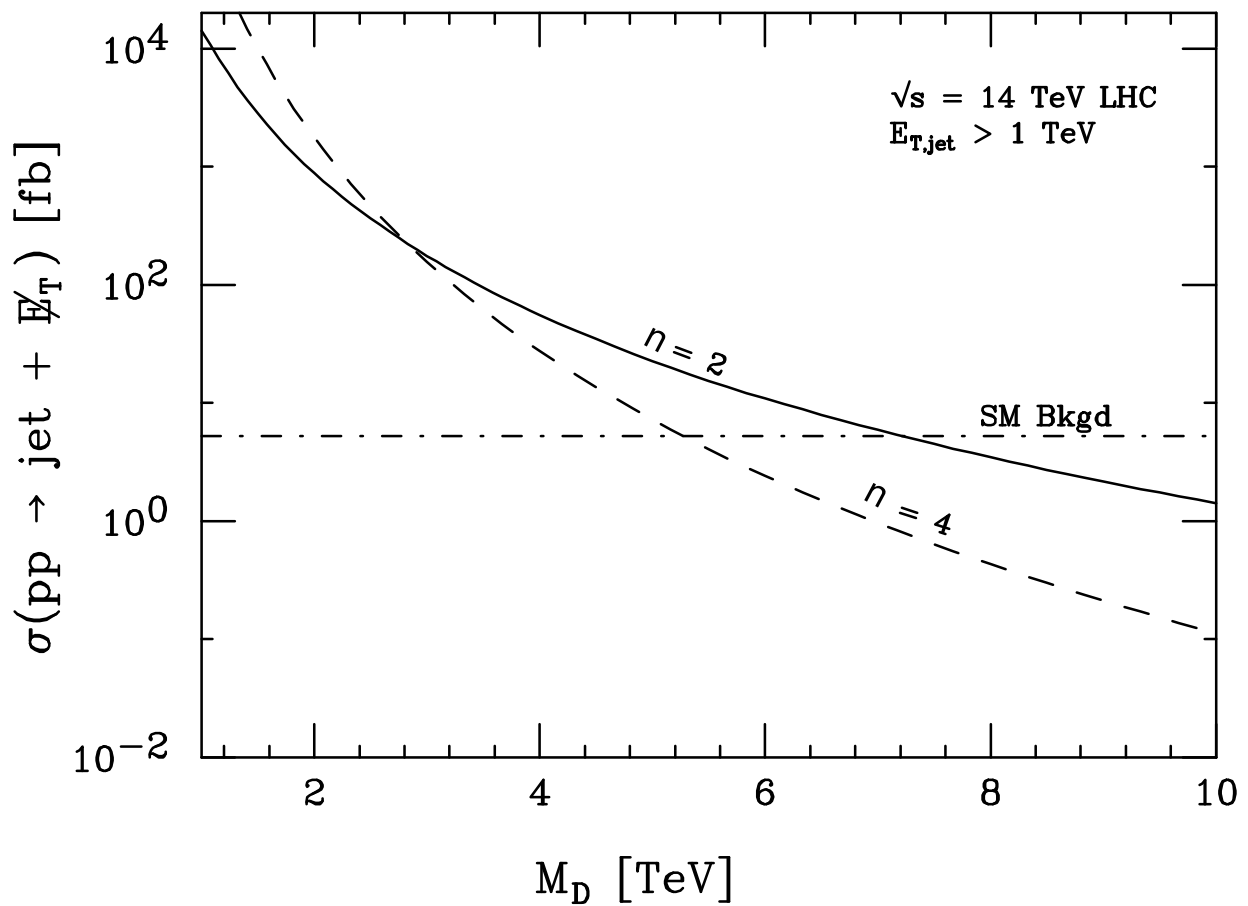
⇒ high energy: huge number of particles produced

LHC: 10^{30} massive gravitons of intensity 10^{-30} each



Signal: missing energy

Angular distribution ⇒ spin of the graviton



no observation \Rightarrow

$$R_{\perp} \lesssim 10^{-2} - 10^{-12} \text{ mm } (n = 2 - 6); 95\% \text{ CL}$$

- more dimensions \Rightarrow weaker limits

Limits on R_{\perp} in mm

Experiment	$R_{\perp}(n = 2)$	$R_{\perp}(n = 4)$	$R_{\perp}(n = 6)$
Collider bounds			
LEP 2	4.8×10^{-1}	1.9×10^{-8}	6.8×10^{-11}
Tevatron	5.5×10^{-1}	1.4×10^{-8}	4.1×10^{-11}
LHC	4.5×10^{-3}	5.6×10^{-10}	2.7×10^{-12}
NLC	1.2×10^{-2}	1.2×10^{-9}	6.5×10^{-12}
Astrophysics/cosmology bounds			
SN1987A	3×10^{-4}	1×10^{-8}	6×10^{-10}
COMPTEL	5×10^{-5}	-	-

Large TeV dimensions

longitudinal dimensions: $R^{-1} \lesssim M_s \Rightarrow$

R^{-1} first scale of new physics I.A. '90

increasing the energy

- could happen for some of the internal dims
- explain coupling constant ratios g_2/g_3
- susy breaking
- fermion masses displace light generations

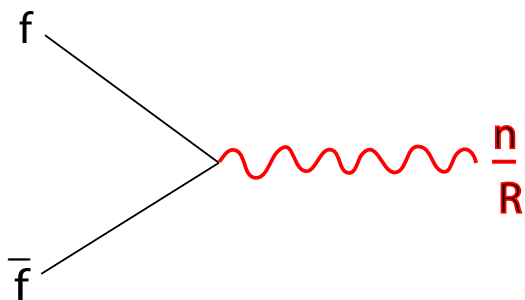
Massive tower of Kaluza Klein modes
for Standard Model particles

$$M_n^2 = M_0^2 + \frac{n^2}{R^2} \quad ; \quad n = \pm 1, \pm 2, \dots$$

\Rightarrow excited states of photon, W^\pm , Z , gluons

Localized fermions (on 3-brane intersections)

⇒ single production of KK modes

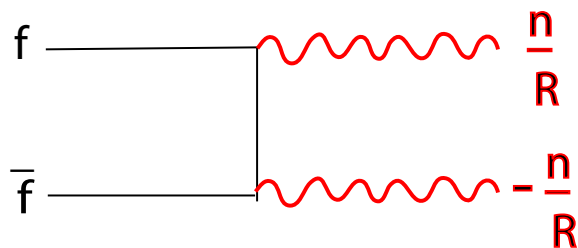


I.A.-Benakli '94

- strong bounds indirect effects: $R^{-1} \gtrsim 3\text{TeV}$
- new resonances but at most $n = 1$

Otherwise KK momentum conservation

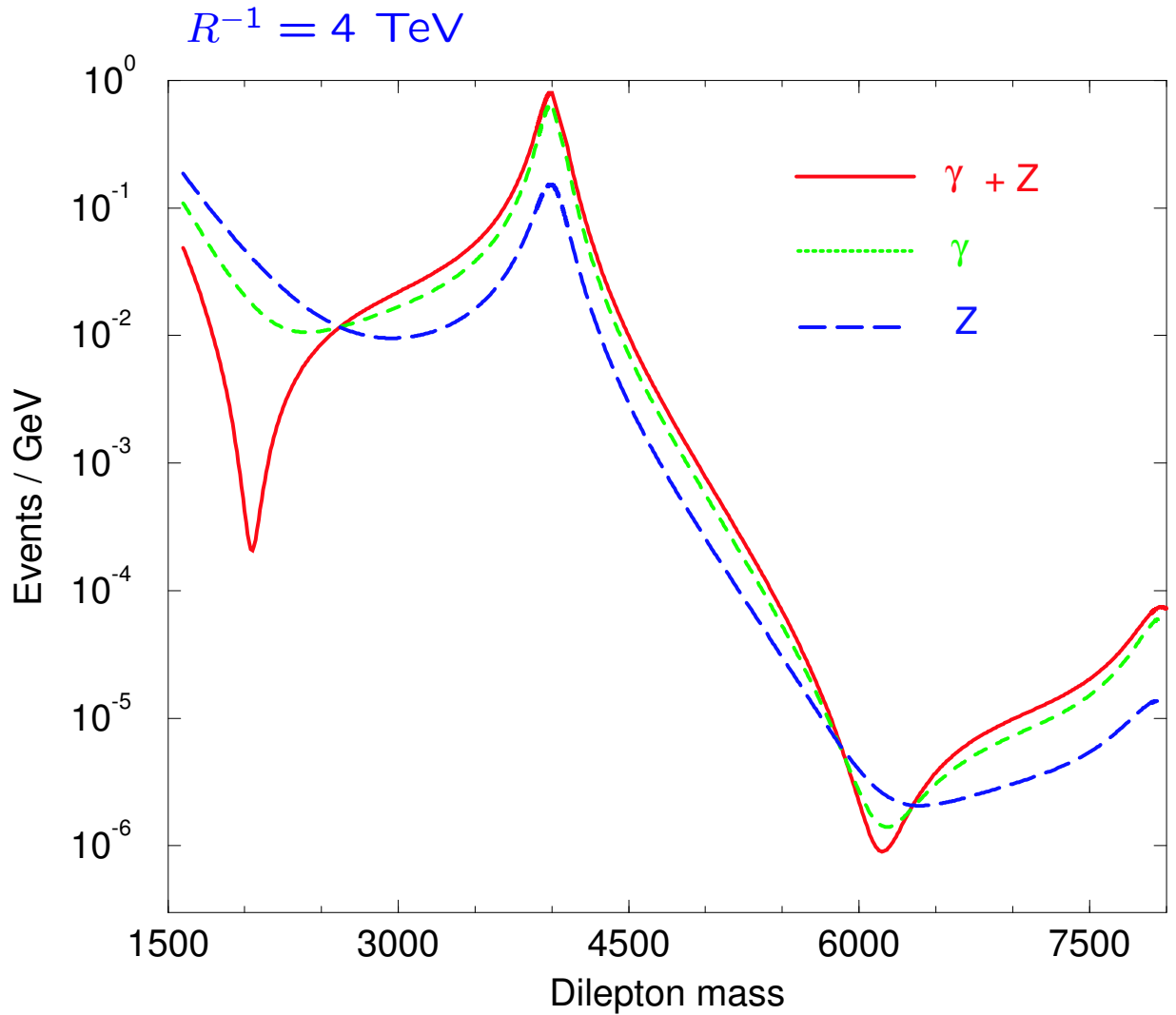
⇒ pair production of KK modes (universal dims)



- weak bounds $R^{-1} \gtrsim 300\text{-}500\text{ GeV}$
- no resonances
- lightest KK stable ⇒ dark matter candidate

Servant-Tait '02

I.A.-Benakli-Quiros '94, '99



- no observation in dijets

$$\Rightarrow R^{-1} \gtrsim 20 \text{ TeV ; 95\% CL}$$

- more than one dimension \Rightarrow stronger limits

Massive string vibrations \Rightarrow indirect effects

virtual exchanges \Rightarrow effective interactions

e.g. four-fermion operators

Actual limits: Matter fermions on

- same set of branes $\Rightarrow M_s \gtrsim 500$ GeV

dim-8: $\frac{g^2}{M_s^4}(\bar{\psi}\partial\psi)^2$ Cullen-Perelstein-Peskin '00

- brane intersections $\Rightarrow M_s \gtrsim 2 - 3$ TeV

dim-6: $\frac{g^2}{M_s^2}(\bar{\psi}\psi)^2$ I.A.-Benakli-Laugier '00

High energies \Rightarrow

- direct production: string physics

- strong gravity: production of micro-black holes?

Giddings-Thomas, Dimopoulos-Landsberg '01

- global SUSY:

- No need to be there **at least for hierarchy**
- New ways of breaking

using extra dimensions

branes at angles/internal magnetic fields

- SUGRA: probably unbroken in the bulk \Rightarrow
very weakly broken

- New forces at submm scales
e.g. radion, graviphoton
- Non linear realization on branes
SM + (light) goldstino

Energy density: $\Lambda_{\text{bulk}}, \Lambda_{\text{brane}}$

generic non-SUSY string model \Rightarrow

$$\Lambda_{\text{bulk}} \sim M_s^{4+n} \Rightarrow \Lambda_{\text{brane}} \sim M_s^{4+n} R_{\perp}^n \sim M_s^2 M_P^2$$

analog in softly broken SUSY: $m_{\text{SUSY}}^2 \Lambda_{UV}^2$

quadratic divergence to Λ

vanishing if bulk is (approximately) SUSY

$$\Lambda_{\text{brane}} \sim M_s^4 \Rightarrow \Lambda_{\text{bulk}} \sim M_s^4 / R_{\perp}^n$$

Prediction: possible new forces at submm scales

e.g. radion $\equiv \ln R_{\perp}$

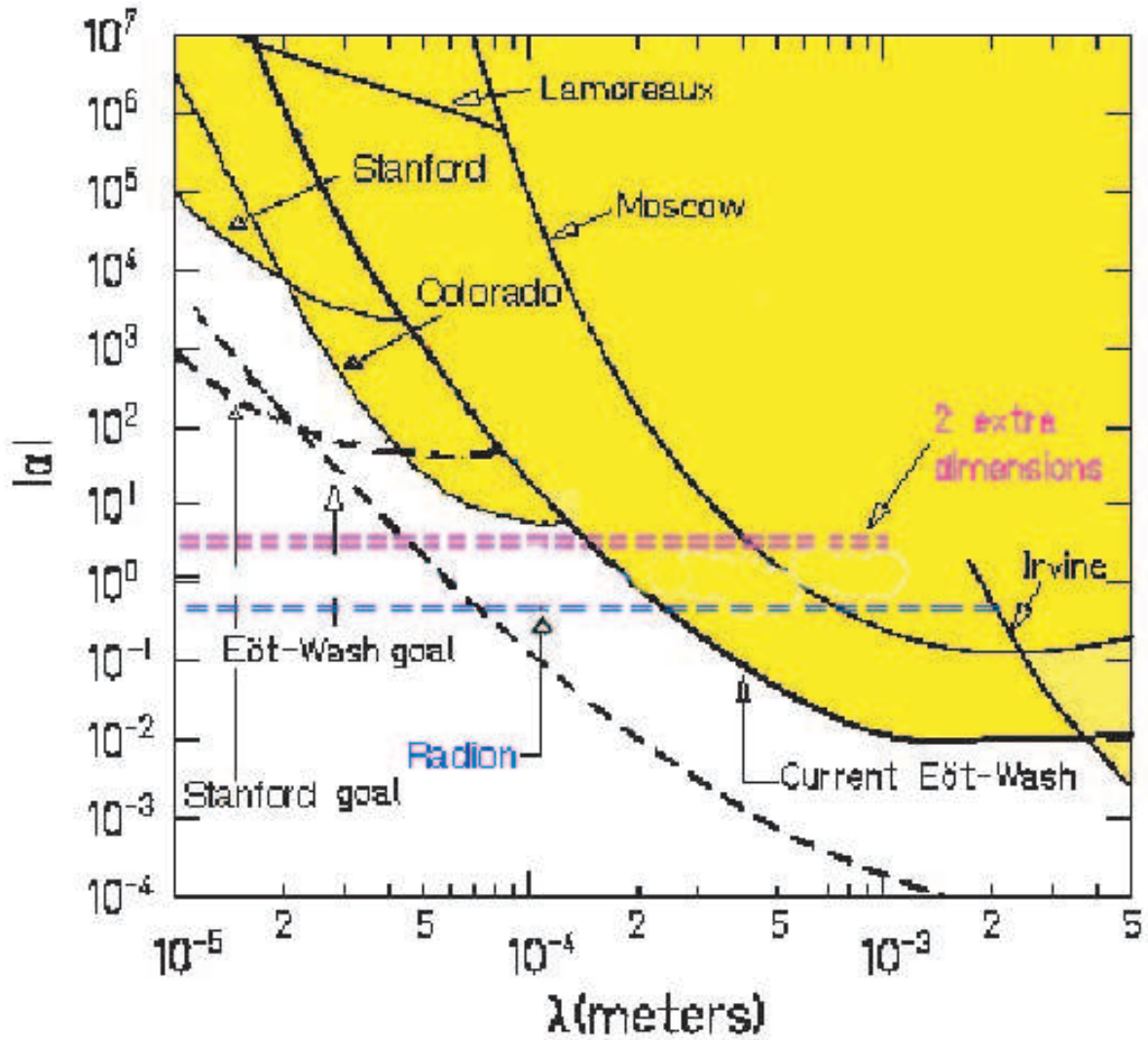
mass: $(\text{TeV})^2 / M_P \sim 10^{-4} \text{ eV} \rightarrow \text{mm range}$

coupling: $\frac{1}{m} \frac{\partial m}{\partial \ln R_{\perp}} = \sqrt{\frac{n}{n+2}} \times \text{gravity}$

\Rightarrow can be experimentally tested for all $n \geq 2$

I.A.-Benakli-Maillard-Laugier '02

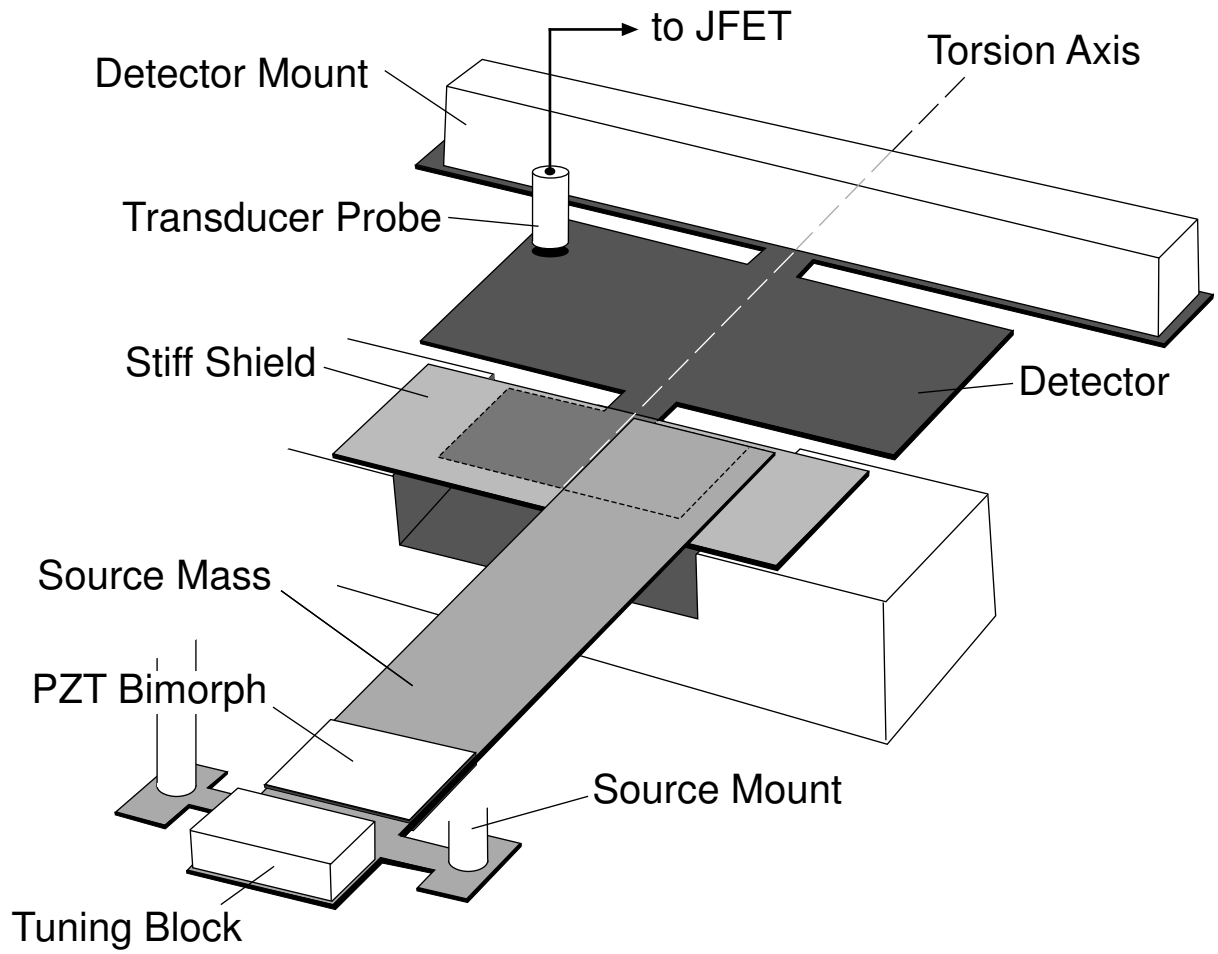
$$V(r) = -G \frac{m_1 m_2}{r} (1 + \alpha e^{-r/\lambda})$$



Radion $\Rightarrow M_* \gtrsim 3 - 4.5 \text{ TeV}$ 95% CL ($n=2-6$)

Adelberger et al. '04

Long-Chan-Churnside-Gulbis-Varney-Price '03



Light $U(1)$ gauge bosons

I.A.-Kiritsis-Rizos '02

$U(1)$ anomalies \Rightarrow Green-Schwarz mechanism

$$\delta A = d\Lambda \quad \Rightarrow \quad \delta a = -M\Lambda$$

↑ gauge field ↑ axion

$$-\frac{1}{4g_A^2} F_A^2 - \frac{1}{2} (da + MA)^2 + \frac{a}{M} k_I^A \text{Tr} F_I \wedge F_I$$

cancel the anomaly ↑

$$\Rightarrow U(1)_A \text{ mass: } m_A = g_A M$$

- a : Poincaré dual of a 2-form

from RR closed string sector $da = *dB_2$

- $U(1)_A$ global symmetry remains

(in perturbation)

ex. Baryon and Lepton number needed to
protect proton decay and neutrino masses

$$m_A = g_A M$$

small mass \Rightarrow small coupling

\Rightarrow A in the bulk and a on the brane:

localized mass

$$g_A \sim 1/\sqrt{V_\perp}$$

$$\Rightarrow m_A \gtrsim M_s^2/M_P \simeq 10^{-4} \text{ eV}$$

A propagates in part of the bulk

\Rightarrow new submm forces

$$g_A \sim 1/\sqrt{V_\perp} \gtrsim M_s/M_P \sim 10^{-16}$$

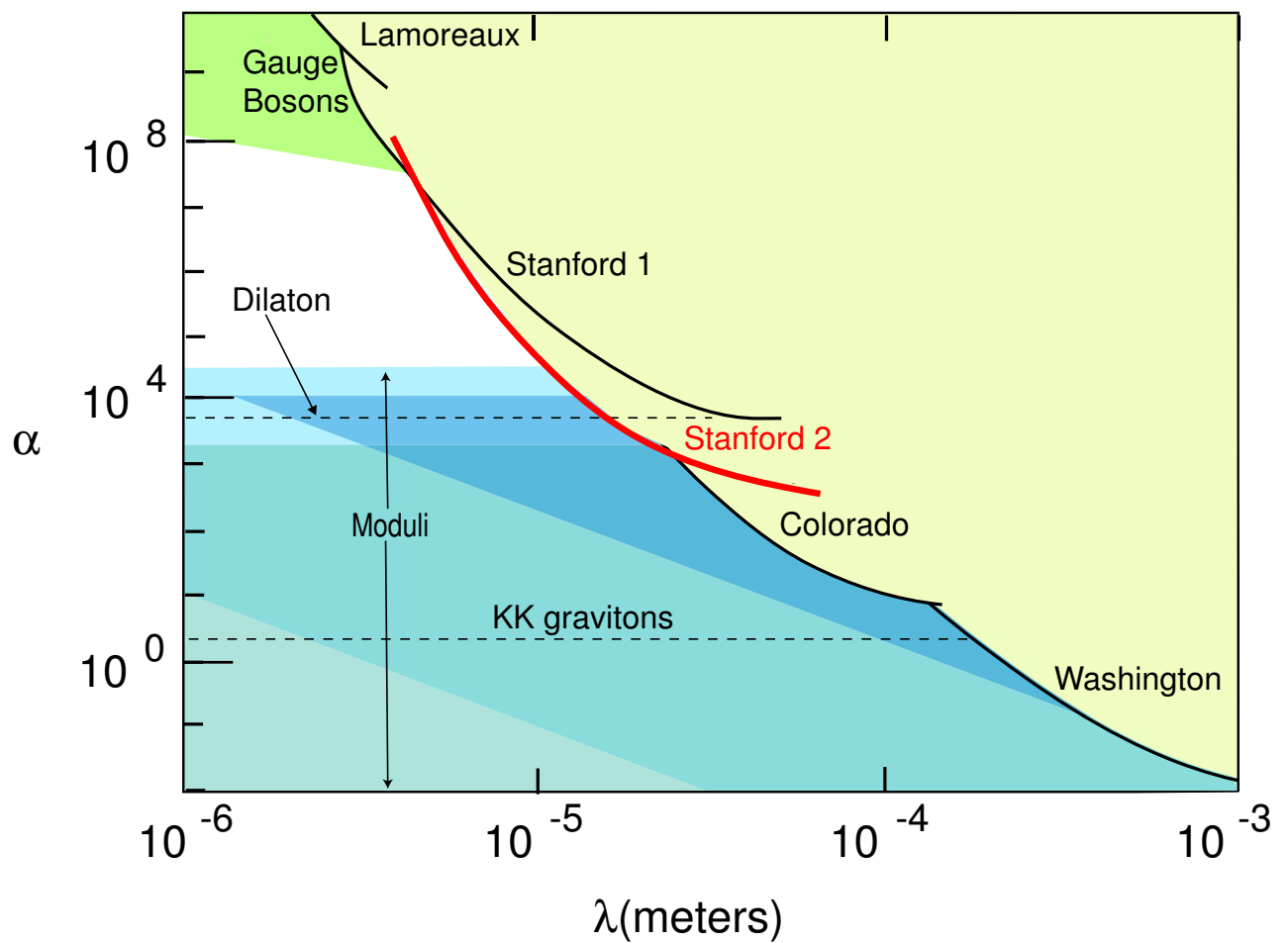
$\Rightarrow \gtrsim 10^6 - 10^8 \times$ gravity

m_{proton}/M_P

supernova \Rightarrow dim of the bulk ≥ 4

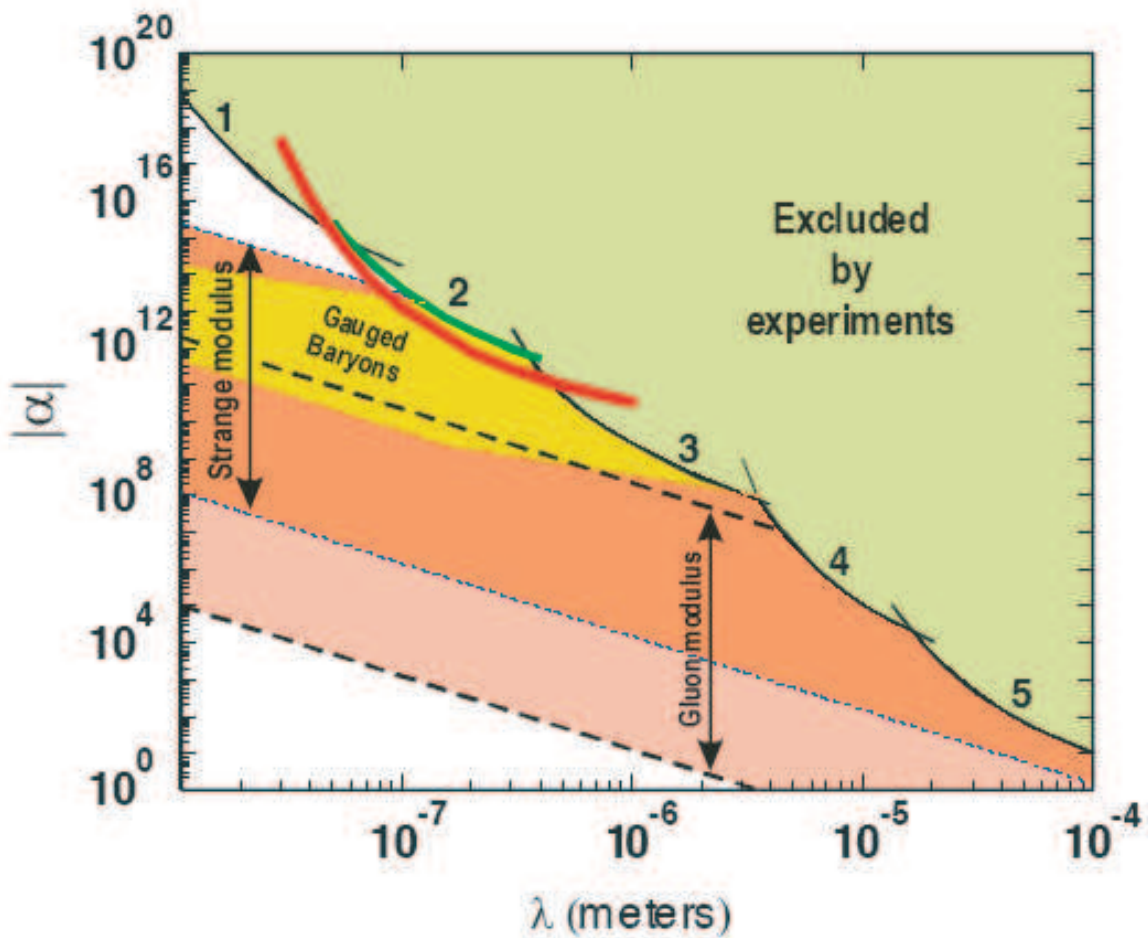
an order of magnitude improvement
on bounds in the range 6-20 μm

Smullin-Geraci-Weld-Chiaverini-Holmes-Kapitulnik '05



an order of magnitude improvement
on bounds in the range 200 nm

Decca-López-Chan-Fischbach-Krause-Jamell '05



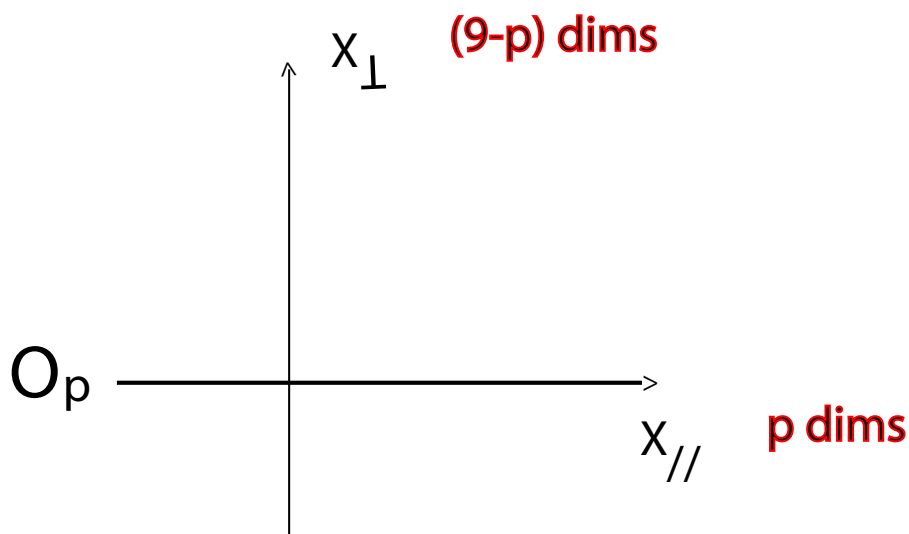
5: Colorado

4: Stanford

3: Lamoureux

1: Mohideen et al.

Orientifold: (hyper)surface where closed strings
change orientation



$$X_{\perp} \rightarrow -X_{\perp} \quad p\text{-plane localized at } X_{\perp} = 0$$

$$z \rightarrow \bar{z} \quad \text{worldsheet orientation flip}$$

non-dynamical object with RR charge \Rightarrow

can have negative tension

Brane supersymmetry breaking

I.A.-Dudas-Sagnotti, Aldazabal-Uranga '99

Stable configurations of branes with orientifolds

- absence of tachyons
- bulk susy breaking suppressed by R_{\perp}

	D	\bar{D}	O	\bar{O}
RR charge	+	-	-	+
tension	+	+	-	-
linear SUSY	Q_e	Q_o	Q_e	Q_o
NL SUSY	Q_o	Q_e		

Model I: DO or $\bar{D}\bar{O}$

local charge conservation, brane SUSY (locally)

Model II: $\bar{D}O$ or $D\bar{O}$

brane SUSY breaking (linear), NL SUSY

Scherk-Schwarz (SS) SUSY breaking

Scherk-Schwarz '79, Rohm '84, Fayet '85

Ferrara-Kounnas-Porrati-Zwirner '88, I.A. '90

Periodicity up to R-symmetry transformation

$$\Phi(y + 2\pi R) = U\Phi(y) \quad U = e^{2\pi i Q} \quad \Rightarrow$$

KK-momentum: $p = \frac{m+Q}{R} \Rightarrow$ mass-shifts

R-symmetry: discrete internal rotation $U^N = 1$

$\Rightarrow Q$ quantized in units of $1/N$

Closed strings: modular invariance \Rightarrow

$$\text{windings } n \rightarrow n, Q \rightarrow Q - n$$

Open strings: $R_{\parallel} \Rightarrow$ like in field theory

$R_{\perp} \Rightarrow$ brane supersymmetry

I.A.-Dudas-Sagnotti '98

Example: $I = S^1/\mathbb{Z}_2$ with SS SUSY breaking

$$O8 \xrightarrow{\pi R} \bar{O}8$$

RR charge: -8

+8

- SS SUSY breaking: 16 D9 branes along I

$\Rightarrow SO(32)$ with fermion mass-shifts

- Model I: 8 D8 branes on $O8$

8 $\bar{D}8$ branes on $\bar{O}8$

$\Rightarrow SO(16) \times SO(16)$ 'SUSY'

- Model II: 8 $\bar{D}8$ branes on $O8$

8 D8 branes on $\bar{O}8$

$\Rightarrow SO(16) \times SO(16)$ with fermions in the sym

$(136, 1) + (1, 136)$

$136 = 135 + 1 \leftarrow$ goldstino

Model III: D away from O

L + NL SUSY

partial breaking " $N = 2 \rightarrow N = 1$ "

8 D8 and 8 $\bar{D}8$ branes in the bulk

O8 $\xrightarrow{\pi R}$ $\bar{O}8$

RR charge: -8

+8

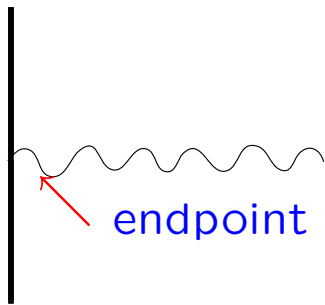
$\Rightarrow U(8) \times U(8)$

$U(1)$: goldstino multiplet

Generic spectrum

N coincident branes $\Rightarrow U(N)$

a-stack



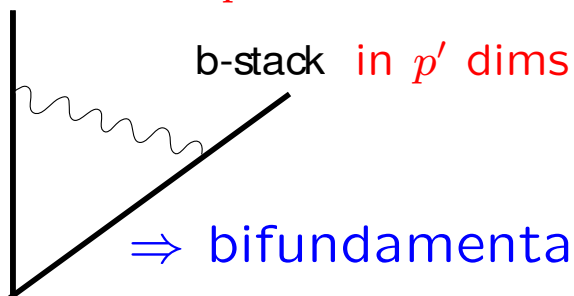
endpoint transformation: N_a or \bar{N}_a

$U(1)_a$ charge: $+1$ or -1

$U(1)$: “baryon” number

- open strings from the same stack \Rightarrow
adjoint gauge multiplets of $U(N_a)$
- stretched between two stacks

a-stack in p dims



\Rightarrow bifundamentals of $U(N_a) \times U(N_b)$

in $p \cap p'$ dims

A D-brane embedding of the Standard Model

I.A.-Kiritsis-Tomaras '00

I.A.-Kiritsis-Rizos-Tomaras '02

- oriented strings \Rightarrow

need at least 4 brane-stacks

- existence of bulk with large dimensions \Rightarrow

minimal choice: $U(3) \times U(2) \times U(1) \times U(1)_{bulk}$


color branes (g_3)


weak branes (g_2)

- also for non-oriented strings

with Baryon and Lepton number symmetries

fermion generation $U(3) \times U(2) \times U(1)$

$$\begin{array}{ll}
 Q & (\mathbf{3}, \mathbf{2}; \mathbf{1}, w, 0)_{1/6} \quad w = \pm 1 \\
 u^c & (\bar{\mathbf{3}}, \mathbf{1}; -\mathbf{1}, 0, x)_{-2/3} \quad x = \pm 1, 0 \\
 d^c & (\bar{\mathbf{3}}, \mathbf{1}; -\mathbf{1}, 0, y)_{1/3} \quad y = \pm 1, 0 \\
 L & (\mathbf{1}, \mathbf{2}; 0, \mathbf{1}, z)_{-1/2} \quad z = \pm 1, 0 \\
 l^c & (\mathbf{1}, \mathbf{1}; 0, 0, 1)_1
 \end{array}$$

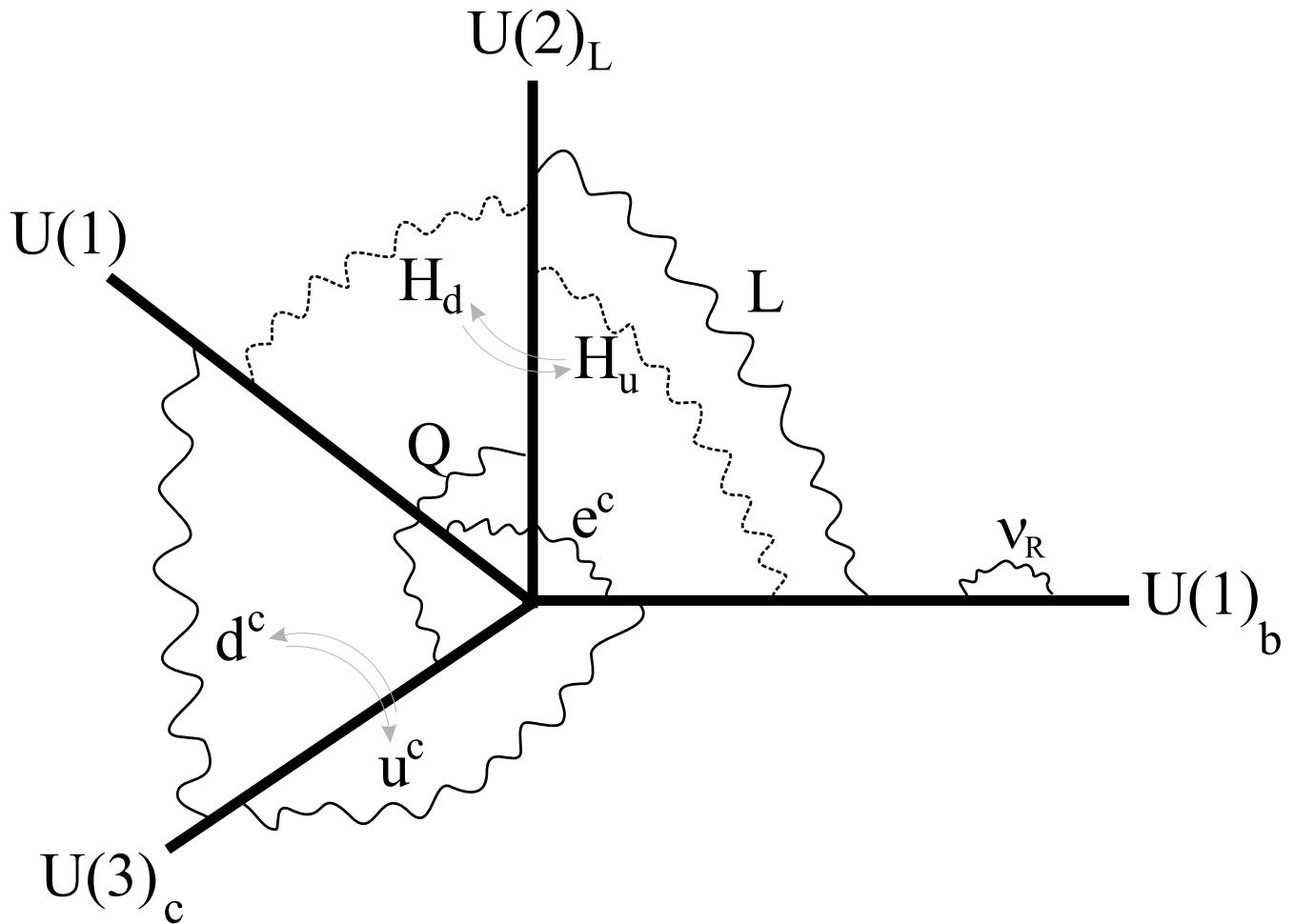
hypercharge $Y = c_1 Q_1 + c_2 Q_2 + c_3 Q_3$

\Rightarrow 4 possibilities:

$$c_3 = -1/3 \quad c_2 = \pm 1/2 \quad x = -1 \quad y = 0 \quad w = \pm 1 \quad z = -1/0$$

$$c_3 = 2/3 \quad c_2 = \pm 1/2 \quad x = 0 \quad y = 1 \quad w = \mp 1 \quad z = -1/0$$

Standard Model on D-branes



- $g_2^2/g_3^2 = R/l_s \Rightarrow$ KK modes for $SU(2)_L$
- $U(1)^4 \Rightarrow$ hypercharge + B, L, PQ global
- $U(1)$ on top of $U(2)$ or $U(3) \Rightarrow$ prediction for $\sin^2 \theta_W$
- ν_R in the bulk \Rightarrow small neutrino masses

The remaining three $U(1)$'s : anomalous

Green-Schwarz anomaly cancellation \Rightarrow

- they become massive (absorb three axions)
- the global symmetries remain in perturbation
- Baryon number \Rightarrow proton stability
- Lepton number \Rightarrow protect small neutrino masses

no Lepton number $\Rightarrow \frac{1}{M_s} LLHH$

\Rightarrow Majorana mass: $\frac{\langle H \rangle^2}{M_s} LL$

\sim GeV

- PQ-type symmetry \Rightarrow electroweak axion

can be explicitly broken by moving slightly away from the orbifold point

$$Y = c_1 Q_1 + c_2 Q_2 + c_3 Q_3 \Rightarrow$$

$$\frac{1}{g_Y^2} = \frac{2c_1^2}{g_1^2} + \frac{4c_2^2}{g_2^2} + \frac{9c_3^2}{g_3^2}$$

$$\begin{aligned} \sin^2 \theta_W &= \frac{g_Y^2}{g_2^2 + g_Y^2} = \frac{1}{g_2^2/g_Y^2 + 1} \\ &= \frac{1}{1 + 4c_2 + 2c_1^2 g_2^2/g_1^2 + 6c_3^2 g_2^2/g_3^2} \end{aligned}$$

fermion generation $U(3) \times U(2) \times U(1)$

$$\begin{array}{ll}
 Q & (\mathbf{3}, \mathbf{2}; 1, w, 0)_{1/6} \quad w = \pm 1 \\
 u^c & (\bar{\mathbf{3}}, \mathbf{1}; -1, 0, x)_{-2/3} \quad x = \pm 1, 0 \\
 d^c & (\bar{\mathbf{3}}, \mathbf{1}; -1, 0, y)_{1/3} \quad y = \pm 1, 0 \\
 L & (\mathbf{1}, \mathbf{2}; 0, 1, z)_{-1/2} \quad z = \pm 1, 0 \\
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hypercharge $Y = c_1 Q_1 + c_2 Q_2 + c_3 Q_3$

\Rightarrow 4 possibilities:

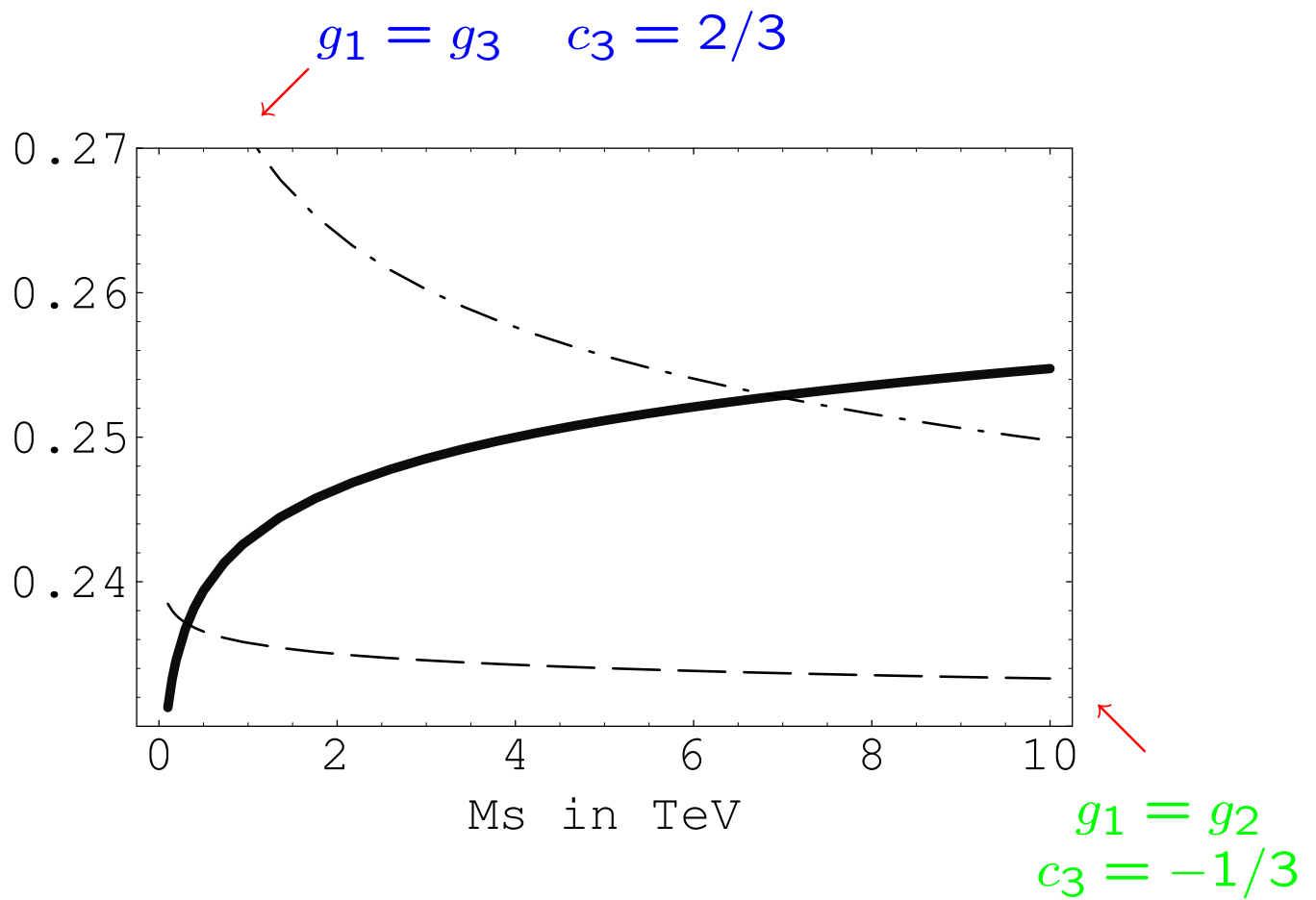
$$c_3 = -1/3 \quad c_2 = \pm 1/2 \quad x = -1 \quad y = 0 \quad w = \pm 1 \quad z = -1/0$$

$$c_3 = 2/3 \quad c_2 = \pm 1/2 \quad x = 0 \quad y = 1 \quad w = \mp 1 \quad z = -1/0$$

$$\sin^2 \theta_W = \frac{1}{2 + 2g_2^2/g_1^2 + 6c_3^2 g_2^2/g_3^2}$$

$$g_1 = g_2 = g_3 \Rightarrow \sin^2 \theta_W = \begin{cases} 3/14 & c_3 = -1/3 \\ 3/20 & c_3 = 2/3 \end{cases}$$

$$\sin^2 \theta_W(M_s)$$



\Rightarrow correct prediction for $\sin^2 \theta_W$
for $M_s \sim$ a few TeV

R-neutrinos: open strings in the bulk $H' L \nu_R$

Arkani Hamed-Dimopoulos-Dvali-March Russell '98

Dienes-Dudas-Gherghetta '98

- $\int d^{4+n}x \bar{\nu} \not{\partial} \nu \quad \nu = (\nu_R, \nu_R^c) \Rightarrow$

$$R_{\perp}^n \int d^4x \sum_m \left\{ \bar{\nu}_{Rm} \not{\partial} \nu_{Rm} + \bar{\nu}_{Rm}^c \not{\partial} \nu_{Rm}^c + \frac{m}{R_{\perp}} \nu_{Rm} \nu_{Rm}^c + c.c. \right\}$$

- $S_{int} = g_s \int d^4x H(x) L(x) \nu_R(x, y=0)$

$$\langle H \rangle = v \Rightarrow \text{mass-terms: } \frac{g_s v}{R_{\perp}^{n/2}} \sum_m \nu_L \nu_{Rm}$$

$$\frac{g_s v}{R_{\perp}^{n/2}} \ll \frac{1}{R_{\perp}} \Leftrightarrow g_s v \ll R_{\perp}^{n/2-1} \text{ in string units} \Rightarrow$$

- $m \neq 0$: masses for KK ν_m unaffected

- $m = 0$: Dirac neutrino masses

$$m_{\nu} \simeq \frac{g_s v}{R_{\perp}^{n/2}} \simeq \frac{g_s}{g^2} v \frac{M_s}{M_p}$$

$$\simeq 10^{-3} - 10^{-2} \text{ eV for } M_s \simeq 1 - 10 \text{ TeV}$$

In principle one $\nu_R \Rightarrow$

both solar and atmospheric oscillations

two frequencies: solar $\leftrightarrow m_\nu \ll$

atmospheric \leftrightarrow 1st KK excitation

however cannot be made realistic

e.g. KK modes \rightarrow important sterile component

\Rightarrow need to introduce three ν_R^i (at least 2)

explain oscillations in the traditional way

- only from zero modes ν_{R0}^i

- make KK modes heavy

Davoudiasl-Langacker-Perelstein '02

Conclusions

TeV strings and large extra dimensions:

Physical reality or imagination?

Well motivated theoretical framework

with many testable experimental predictions

new resonances, missing energy

Stimulus for micro-gravity experiments

look for new forces at short distances

higher dim graviton, scalars, gauge fields

Internal magnetic fields

- Type I string theory compactified in 4d on 6d Calabi-Yau

⇒ $N = 2$ SUSY in the bulk, $N = 1$ on branes

- Magnetic fluxes on 2-cycles

⇒ SUSY breaking

Dirac quantization: $H = \frac{m}{nA} \equiv \frac{p}{A}$

H : constant magnetic field

m : units of magnetic flux

n : brane wrapping

A : area of the 2-cycle

Spin-dependent mass shifts for all charged states

$$[p_i, p_j] = iqH\epsilon_{ij} \quad q: \text{charge}$$

⇒ Landau spectrum

6d \rightarrow 4d on T^2 with abelian magnetic field H

$$\delta M^2 = (2k + 1)|qH| + 2qH \cdot \Sigma \leftarrow \text{spin operator}$$

$k = 0, 1, 2, \dots$: Landau level

Landau multiplicity: mn

• spin-0: $\Sigma = 0 \Rightarrow$ mass gap

• spin-1/2: $\Sigma = \pm 1/2 \Rightarrow$ chiral 0-mode

$$k = 0 \quad : \quad \delta M^2 = |qH| \pm qH$$

$$\Rightarrow \delta M^2 = 0 \quad \text{for } \Sigma = -1/2 \text{ (} qH > 0 \text{)}$$

• spin-1: $\Sigma = \pm 1 \Rightarrow$ tachyon

Nielsen-Olesen instability

$$k = 0 \quad : \quad \delta M^2 = |qH| \pm 2qH$$

$$\Rightarrow \delta M^2 = -qH \quad \text{for } \Sigma = -1 \text{ (} qH > 0 \text{)}$$

Exact open string description:

$$q \rightarrow q_L + q_R \quad \text{endpoint charges}$$

$$qH \rightarrow \theta_L + \theta_R \quad ; \quad \theta_{L,R} = \arctan q_{L,R} H \alpha'$$

weak field limit \Rightarrow field theory

$$H \text{ constant} \Rightarrow F_{kl} = \epsilon_{kl} H \quad A_k = -\frac{1}{2} F_{kl} x^l$$

world-sheet boundary action:

$$q \int A_k \partial x^k = -H \int \left(q_L x^k \overleftrightarrow{\partial} x^l \Big|_{\sigma=0} + q_R x^k \overleftrightarrow{\partial} x^l \Big|_{\sigma=\pi} \right)$$

internal rotation current

$$\Rightarrow \text{frequency shift by } \theta_{L,R} : \tan \theta_{L,R} = q_{L,R} H$$

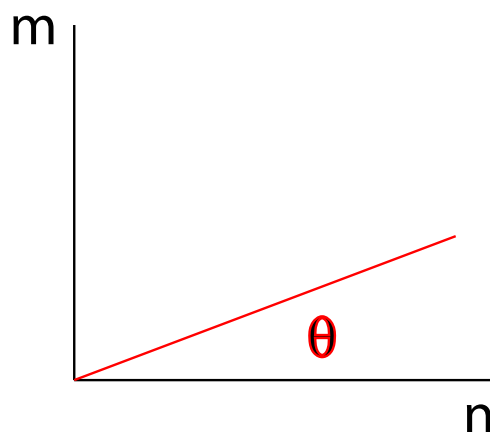
T-dual representation: branes at angles
magnetized D9-brane wrapped on T^2

$$H = \frac{m}{n} \frac{1}{R_1 R_2}$$

T-duality: $R_2 \rightarrow \alpha'/R_2 \equiv \tilde{R}_2 \Rightarrow$ D8-brane
wrapped around a direction of angle θ in T^2

$$H = \frac{m}{n} \frac{\tilde{R}_2}{R_1} = \tan \theta$$

(m, n) : wrapping numbers around (\tilde{R}_2, R_1)

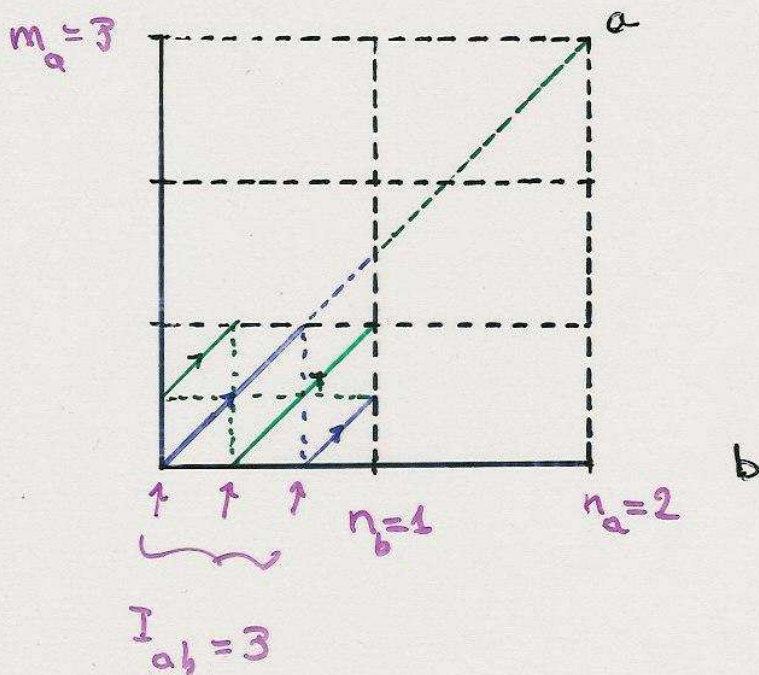


Chirality = intersection number

e.g. $I_{ab} = m_a n_b - m_b n_a$

= intersection nb of branes a, b

ex. $m_b = 0$ $n_b = 1$ $\Rightarrow I_{ab} = m_a$



$(T^2)^3$ generalization: H_I with $I = 1, 2, 3$

$$\delta M^2 = \sum_I \{(2k_I + 1)|qH_I| + 2qH_I\Sigma_I\}$$

- spin-1/2: one chiral 0-mode

$$\delta M^2 = 0 \text{ for } k_I = 0 \text{ and } \Sigma_I = -1/2 \text{ (} qH_I > 0 \text{)}$$

- spin-1: tachyon can be avoided Bachas 95

$$\begin{array}{r} |H_1| + |H_2| - |H_3| > 0 \\ |H_1| - |H_2| + |H_3| > 0 \\ - |H_1| + |H_2| + |H_3| > 0 \end{array}$$

massless scalar \Leftrightarrow partial brane susy restoration

Angelantonj-I.A.-Dudas-Sagnotti 00

$$\theta_1 + \theta_2 + \theta_3 = 0$$

Generic spectrum

Turn on H_I^a in several $U(1)_a$ directions

\Rightarrow Gauge group: $\prod_a U(N_a) \leftarrow SU(N_a) \times U(1)_a$

- Neutral strings: adjoint representations

\Rightarrow massless gauge supermultiplets

- Charged strings \Rightarrow massless chiral fermions

but in general massive scalars

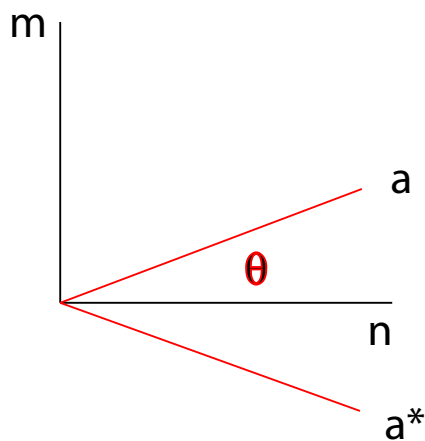
\Rightarrow Generic spectrum of split SUSY:

- massless gauginos
- massive squarks and sleptons
- massless Higgs \Leftrightarrow non chiral susy intersection
two Higgs multiplets

D-brane a : (m, n) ; $n > 0$ anti-brane: $(m, -n)$

Orientifold: $(0, x)$

Mirror brane a^* : $(-m, n)$



multiplicities: nb of intersections in $(1, 1)$

$$(N_a, \bar{N}_b): I_{ab} = \prod_I (m_I^a n_I^b - n_I^a m_I^b)$$

$$(N_a, N_b): I_{ab^*} = \prod_I (m_I^a n_I^b + n_I^a m_I^b)$$

same stack: antisymmetric or symmetric

$$I_{aa^*} = \prod_I \left\{ \frac{1}{2} (2m_I^a n_I^a \mp 2m_I^a) \pm 2m_I^a \right\} = \begin{cases} A: \frac{1}{2} (\prod_I 2m_I^a) (\prod_J n_J^a + 1) \\ S: \frac{1}{2} (\prod_I 2m_I^a) (\prod_J n_J^a - 1) \end{cases}$$

nb of intersections along $(0, x)$

- **Gauge coupling unification**

I.A.-Dimopoulos '04

- **non-abelian** $\alpha_2 = \alpha_3$ at 1%

guaranteed by:

(i) the correct SM spectrum:

no chiral color sextets, weak triplets,

antiquark doublets

(ii) weak magnetic fields

$\Rightarrow M_{\text{GUT}/\text{comp}} \sim M_s/3$

- **weak angle** $\sin^2 \theta_W = \frac{3}{8}$

possible

e.g. using SM embedding in 3 brane stacks

Minimal Standard Model embedding

New possibilities using intersecting branes

- no large dimensions for low string scale
- no need for B or L conservation
- but need $\sin^2 \theta_W = \frac{3}{8}$

General analysis using 3 brane stacks

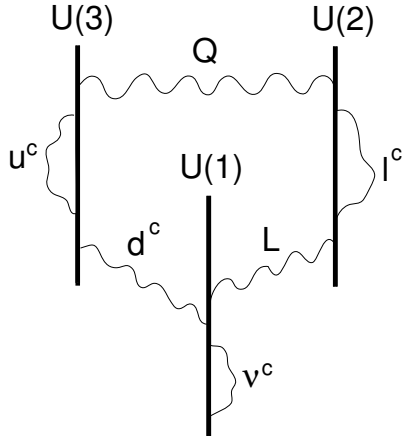
$$\Rightarrow U(3) \times U(2) \times U(1)$$

antiquarks u^c, d^c ($\bar{3}, 1$):

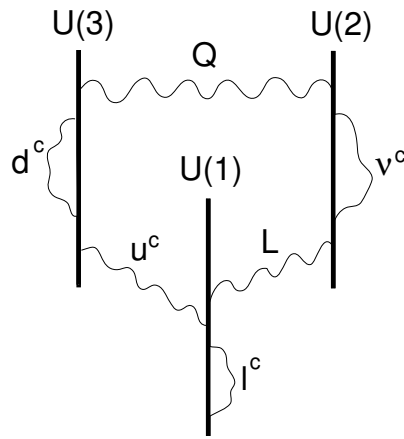
antisymmetric of $U(3)$ or

bifundamental $U(3) \leftrightarrow U(1)$

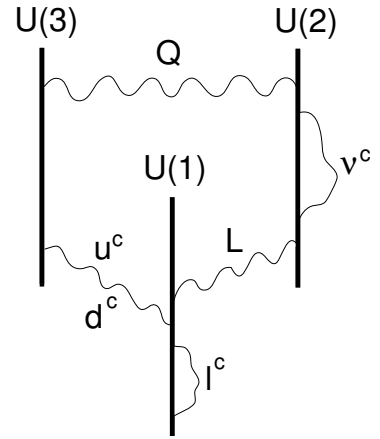
\Rightarrow 3 models: antisymmetric is u^c, d^c or none



Model A



Model B



Model C

Q	$(\mathbf{3}, \mathbf{2}; 1, 1, 0)_{1/6}$	$(\mathbf{3}, \mathbf{2}; 1, \varepsilon_Q, 0)_{1/6}$	$(\mathbf{3}, \mathbf{2}; 1, \varepsilon_Q, 0)_{1/6}$
u^c	$(\bar{\mathbf{3}}, \mathbf{1}; 2, 0, 0)_{-2/3}$	$(\bar{\mathbf{3}}, \mathbf{1}; -1, 0, 1)_{-2/3}$	$(\bar{\mathbf{3}}, \mathbf{1}; -1, 0, 1)_{-2/3}$
d^c	$(\bar{\mathbf{3}}, \mathbf{1}; -1, 0, \varepsilon_d)_{1/3}$	$(\bar{\mathbf{3}}, \mathbf{1}; 2, 0, 0)_{1/3}$	$(\bar{\mathbf{3}}, \mathbf{1}; -1, 0, -1)_{1/3}$
L	$(\mathbf{1}, \mathbf{2}; 0, -1, \varepsilon_L)_{-1/2}$	$(\mathbf{1}, \mathbf{2}; 0, \varepsilon_L, 1)_{-1/2}$	$(\mathbf{1}, \mathbf{2}; 0, \varepsilon_L, 1)_{-1/2}$
l^c	$(\mathbf{1}, \mathbf{1}; 0, 2, 0)_1$	$(\mathbf{1}, \mathbf{1}; 0, 0, -2)_1$	$(\mathbf{1}, \mathbf{1}; 0, 0, -2)_1$
ν^c	$(\mathbf{1}, \mathbf{1}; 0, 0, 2\varepsilon_\nu)_0$	$(\mathbf{1}, \mathbf{1}; 0, 2\varepsilon_\nu, 0)_0$	$(\mathbf{1}, \mathbf{1}; 0, 2\varepsilon_\nu, 0)_0$

$$Y_A = -\frac{1}{3}Q_3 + \frac{1}{2}Q_2$$

$$Y_{B,C} = \frac{1}{6}Q_3 - \frac{1}{2}Q_1$$

$$\text{Model A} \quad : \quad \sin^2 \theta_W = \frac{1}{2 + 2\alpha_2/3\alpha_3} \Big|_{\alpha_2=\alpha_3} = \frac{3}{8}$$

$$\text{Model B, C} \quad : \quad \sin^2 \theta_W = \frac{1}{1 + \alpha_2/2\alpha_1 + \alpha_2/6\alpha_3} \Big|_{\alpha_2=\alpha_3} = \frac{6}{7 + 3\alpha_2/\alpha_1}$$

- Higgs can be easily implemented

massless \Rightarrow susy intersection

$$H_1, H_2 : U(2) \leftrightarrow U(1) \quad \text{like } L$$

Model A

Model B, C

H_1	$(\mathbf{1}, \mathbf{2}; 0, -1, \varepsilon_{H_1})_{-1/2}$	$(\mathbf{1}, \mathbf{2}; 0, \varepsilon_{H_1}, \mathbf{1})_{-1/2}$
H_2	$(\mathbf{1}, \mathbf{2}; 0, \mathbf{1}, \varepsilon_{H_2})_{1/2}$	$(\mathbf{1}, \mathbf{2}; 0, \varepsilon_{H_2}, -\mathbf{1})_{1/2}$

- 2 extra $U(1)$'s
 - Model A,B: one combination can be $B - L$
broken by a SM singlet VEV at high scale
or survive at low energies
 - Model C: Baryon symmetry
 - The other/both is/are anomalous

Gaugino masses: protected by R-symmetry

but broken in 4d SUGRA by the gravitino mass

Two possible ways for generating $m_{1/2}$:

(1) via gravity (brane susy) \Rightarrow

generate $m_{1/2}$ from $m_{3/2}$

one gravitational loop: 1 handle + 1 boundary

$$\Rightarrow m_{1/2} \sim g_s^2 \frac{m_{3/2}^3}{M_s^2}$$

I.A.-Taylor '04

(2) keep gravity subdominant \Rightarrow

generate $m_{1/2}$ from brane α' -corrections

two gauge loops: 3 boundaries

$$\Rightarrow m_{1/2} \sim g_s^2 \frac{m_0^4}{M_s^3}$$

I.A.-Narain-Taylor '05

gauginos: open strings

\Rightarrow at least one boundary (brane) $h \geq 1$

$N = 2$ superconformal charge:

$3/2$ units for each (chiral) gaugino

± 1 unit for each 2d supercurrent insertion T_F

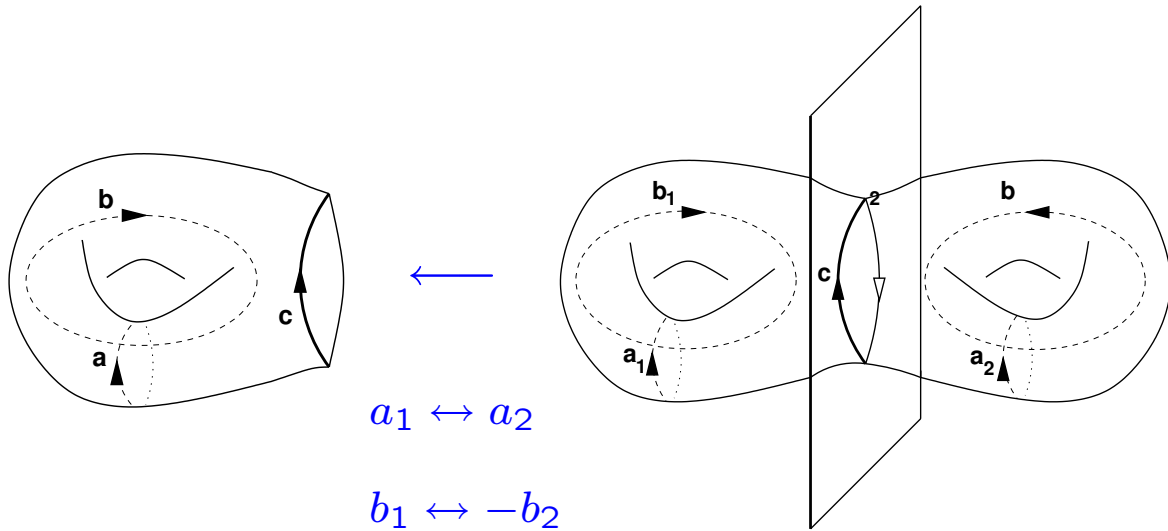
\Rightarrow at least 3 T_F insertions

lowest order (effective genus): $g + h/2 = 3/2$

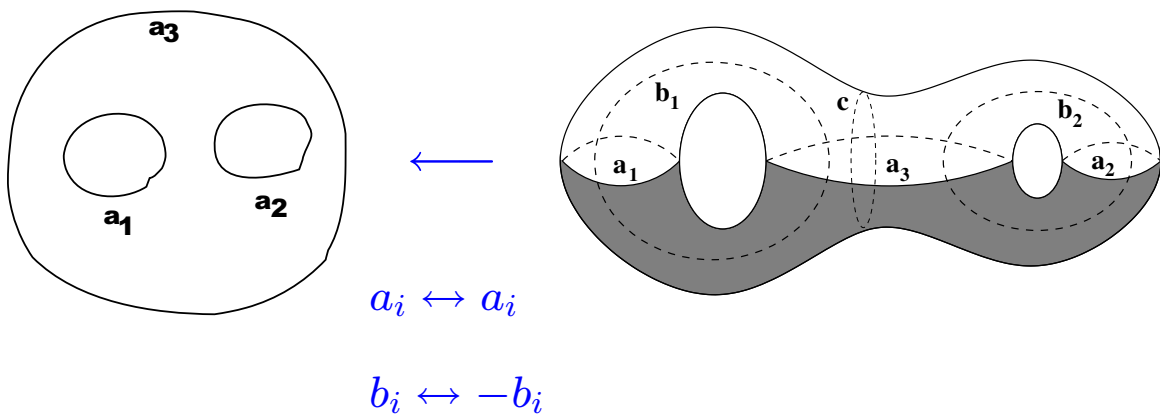
independently of the source of SUSY breaking!


Oriented case

(1) $g = 1 \quad h = 1$ from mirror involution of $g = 2$



(1) $g = 0 \quad h = 3$ from mirror involution of $g = 2$



Topological partition function F_g  genus g
computes $N = 2$ SUSY F-terms

AGNT, BCOV '93

$$F_g \int d^4\theta W_{N=2}^{2g} \rightarrow F_g R^2 T^{2g-2}$$

F_g : moduli dependent function

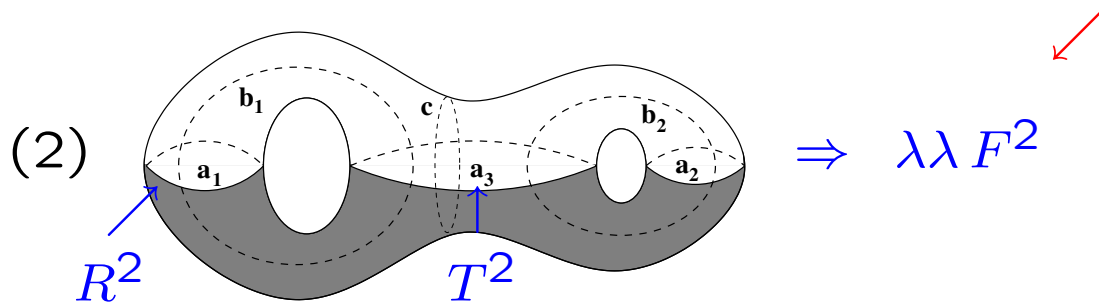
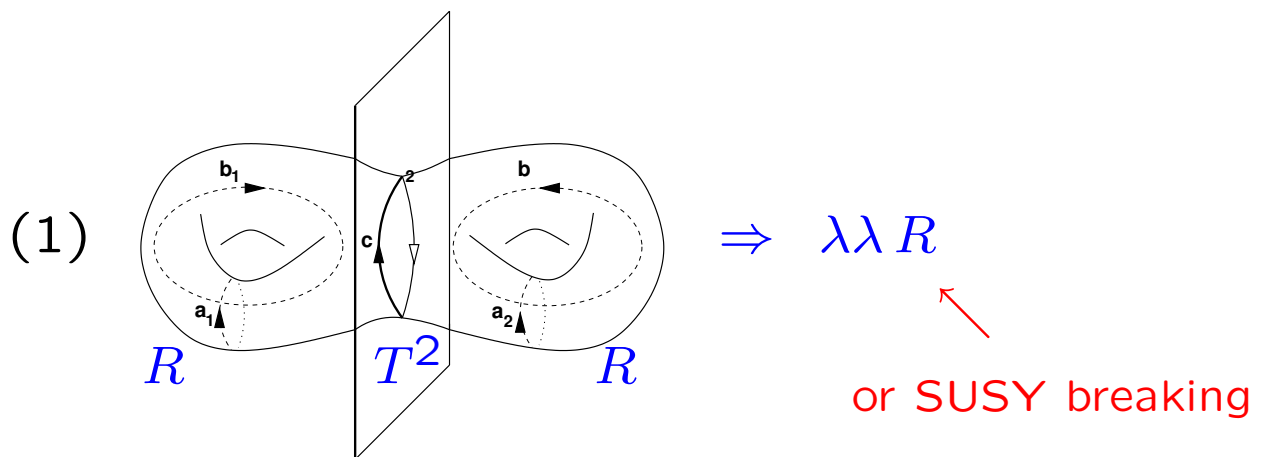
Weyl superfield: $W_{N=2} = T + \theta^2 R + \dots$

T : graviphoton field strength

R : Riemann tensor

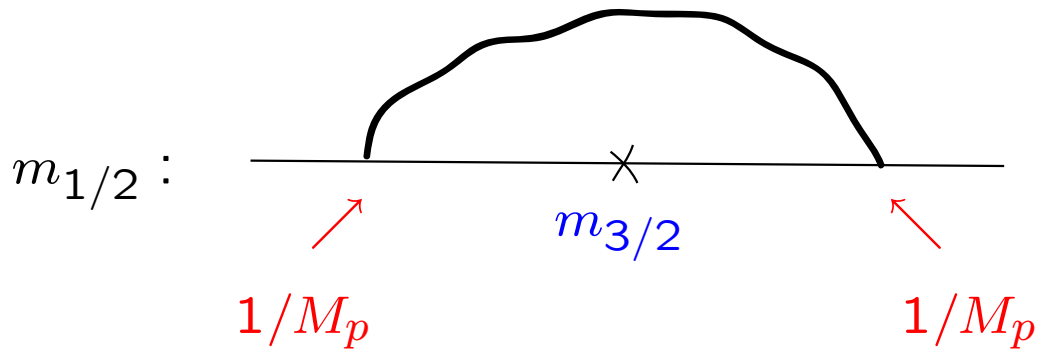
$$F_2 \int d^4\theta W_{N=2}^4 \rightarrow F_2 R^2 T^2$$

- graviphoton vertex $T = (\text{gaugino})^2$
- graviton vertex = (gauge field)²

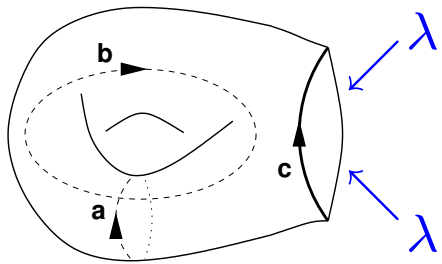


SUSY breaking: $R \rightarrow \langle \text{gravity auxiliary field} \rangle$

$F \rightarrow \langle D \rangle$



$$\sim \frac{m_{3/2}}{M_p^2} \times \begin{cases} \Lambda_{UV}^2 & \text{if quadr. divergent} \\ m_{3/2}^2 & \text{if convergent} \end{cases}$$



$$\sim g_s^2 \frac{m_{3/2}^3}{M_s^2} \quad g_s \sim g^2$$

but it vanishes for orbifolds

I.A.-Taylor '04

- anomaly mediation:

$$m_{1/2} \sim g^2 m_{3/2} \quad g^2 \sim g_s$$

- power of g_s does not match

one loop correction always vanishes

by $N = 2$ superconformal charge

- two loops behave $\sim m_{3/2}^3$

- hierarchy between gaugino and scalar masses

however numerics not very good

unless every loop factor $\sim 10^{-2}$

Sherk-Schwarz along an interval \perp branes

$$\Rightarrow m_{3/2} \sim 1/R$$

$$\text{gravity strength} \Rightarrow R^{-1} = \frac{2}{\alpha_G^2} \frac{M_s^3}{M_p^2} \sim 10^{13} \text{ GeV}$$

$$\text{for } M_s \sim M_{\text{GUT}} \sim 10^{16} \text{ GeV}$$

$$\bullet m_{1/2} \sim g_s^2 \frac{m_{3/2}^3}{M_s^2} \sim 1 \text{ TeV}$$

$$\text{if every loop-factor} \sim 10^{-2}$$

$$\bullet m_0 \gtrsim g_s \frac{m_{3/2}^2}{M_s} \sim 10^8 \text{ GeV}$$

scalar masses induced at one loop

\Rightarrow split supersymmetry framework

heavy scalars, light fermions

Arkani Hamed-Dimopoulos, Giudice-Romanino '04

SUSY breaking by internal magnetic fields
or equivalently branes at angles

Effective QFT description: D-breaking

magnetic field $H \sim \langle D \rangle$ -term of $U(1)$

$$\langle D \rangle \sim m_0^2$$


 $U(N)$ brane stack

R-symmetry broken by string corrections

\Rightarrow higher-dim effective operators:

I.A.-Narain-Taylor '05

$$F_{(0,3)} \int d^2\theta \mathcal{W}^2 \text{Tr} W^2$$

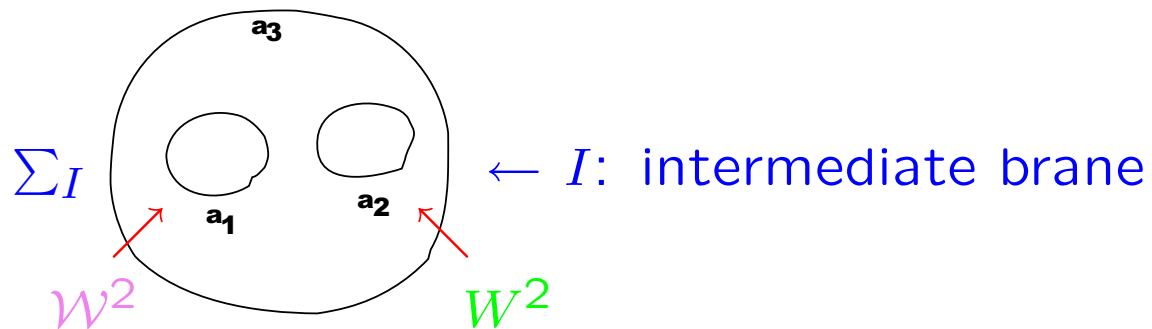
$$\langle \mathcal{W} \rangle = \theta \langle D \rangle$$

$$\Rightarrow m_{1/2} \sim \epsilon^2 \frac{m_0^4}{M_s^3}$$

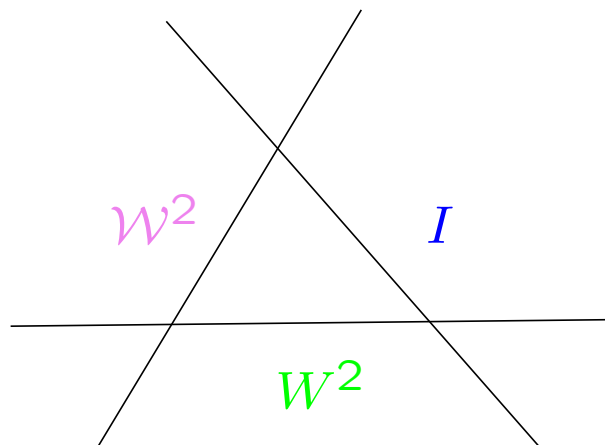
ϵ^2 : 2-loop factor

$$\sim \text{TeV for } m_0 \sim 10^{13} - 10^{14} \text{ GeV}$$

World-sheet with 3 boundaries (2 loops)



T-duality \Rightarrow



$\neq 0$: I -brane away from the intersection
of the other two

- as gauge mediation with string scale gaugino masses

- Higgsino mass

$$\int d^2\theta \mathcal{W}^2 \bar{D}^2 \bar{H}_1 \bar{H}_2 \Rightarrow \mu \sim \epsilon \frac{m_0^4}{M_s^3} \lesssim m_{1/2}$$

\nearrow
 $\psi_1 \psi_2$

- Simple toroidal models

gauge multiplets: $N = 4$ (or $N = 2$) SUSY

\Rightarrow Dirac gaugino masses without \mathbb{R}

$$\int d^2\theta \mathcal{W} \text{Tr} W A \Rightarrow m_D \sim \epsilon \frac{m_0^2}{M_s} \quad \text{1-loop factor}$$

$N = 2$ vector = $N = 1$ vector W + chiral A

they can still be consistent with unification

in intermediate energy scales $\sim 10^7 - 10^{13}$ GeV

I.A.-Benakli-Delgado-Quirós-Tuckmantel '05

Conclusions

Gaugino masses from string loops:

High string scale \Rightarrow hierarchy $m_0 \gg m_{1/2}$

1) Majorana masses

- gravity 'mediation' $\Rightarrow m_{1/2}^2 \sim m_0^3/M_s$
- gauge 'mediation' $\Rightarrow m_{1/2} \sim m_0^4/M_s^3$

2) Dirac masses $\Rightarrow m_D \sim m_0^2/M_s$

evading the hierarchy:

$M_s \rightarrow M_{\text{hyp}}, m_0^2 \rightarrow D$ in a SUSY sector

$m_0^{\text{SM}} \sim m_D$ from 2-loops